## AQA

Please write clearly in block capitals.

Centre number

|  |  |  |  |  |
| :--- | :--- | :--- | :--- | :--- |

Candidate number

|  |  |  |  |
| :--- | :--- | :--- | :--- |

Surname
Forename(s)
Candidate signature
I declare this is my own work.

## A-level PHYSICS

## Paper 3

## Section B Astrophysics

Thursday 15 June 2023 Morning

## Materials

For this paper you must have:

- a pencil and a ruler
- a scientific calculator
- a Data and Formulae Booklet
- a protractor.


## Instructions

- Use black ink or black ball-point pen.
- Fill in the boxes at the top of this page.
- Answer all questions.
- You must answer the questions in the spaces provided. Do not write outside the box around each page or on blank pages.
- If you need extra space for your answer(s), use the lined pages at the end of this book. Write the question number against your answer(s).
- Do all rough work in this book. Cross through any work you do not want to be marked.

Time allowed: The total time for both sections of this paper is 2 hours. You are advised to spend approximately 50 minutes on this section.

- Show all your working.


## Information

- The marks for questions are shown in brackets.
- The maximum mark for this paper is 35 .
- You are expected to use a scientific calculator where appropriate.
- A Data and Formulae Booklet is provided as a loose insert.


## Section B

Answer all questions in this section.

| $\mathbf{0}$ | $\mathbf{1}$ | .1 |
| :--- | :--- | :--- | Draw a labelled diagram to define the parsec (pc).

Table 1 shows data for two stars: Rigel and the Sun.
Table 1

| Star | Surface temperature / K | Absolute magnitude | Mass / kg |
| :---: | :---: | :---: | :---: |
| Rigel | 12000 | -7.84 | $3.6 \times 10^{31}$ |
| Sun | 5700 | 4.83 | $2.0 \times 10^{30}$ |


| 0 | $\mathbf{1}$. | $\mathbf{2}$ State the spectral class of Rigel. |
| :--- | :--- | :--- |

$\qquad$

| 0 | 1 | 3 | The apparent magnitude of Rigel is 0.11 |
| :--- | :--- | :--- | :--- |

Calculate, in pc, the distance from Rigel to the Earth.

| 0 | 1 | .4 | Figure 1 shows a Hertzsprung-Russell (HR) diagram. |
| :--- | :--- | :--- | :--- |

Figure 1


Draw a line on Figure 1 to show the evolution of the Sun from formation to white dwarf.

| 0 | 1 | $\mathbf{5}$ One stage in the evolution of Rigel includes the emission of a gamma ray burst. $. . .0 \mid$ |
| :--- | :--- | :--- |

Outline the circumstances during which a gamma ray burst will be emitted by Rigel.
$\qquad$
$\qquad$
$\qquad$
$\qquad$
$\qquad$

0. 1.5
[2 marks]
Turn over for the next question Turn over

| $\mathbf{0}$ | $\mathbf{2}$ | $\mathbf{1}$ State what is meant by normal adjustment when applied to an astronomical |
| :--- | :--- | :--- | refracting telescope.

$\qquad$
$\qquad$
$\qquad$

| $\mathbf{0}$ | $\mathbf{2} .2$ Which combination of lenses gives the largest angular magnification when used as |
| :--- | :--- | :--- | an astronomical telescope in normal adjustment?

Tick ( $\checkmark$ ) one box.

| Objective lens |  | Eyepiece lens |  |
| :---: | :---: | :---: | :---: |
| Focal length / cm | Type | Focal length / cm | Type |
| 5 | diverging | 100 | converging |
| 5 | converging | 100 | converging |
| 100 | diverging | 5 | converging |
| 100 | converging | 5 | converging |

V1031 and WASP-82 are two stars in the constellation Orion.
V1031 appears 40 times brighter than WASP-82 when viewed from Earth.
The apparent magnitude of V 1031 is 6.0

| $\mathbf{0}$ | $\mathbf{2} .3$ Calculate the apparent magnitude of WASP-82. |
| :--- | :--- | :--- | :--- |

$\qquad$

| $\mathbf{0}$ | $\mathbf{2} .4$ | V 1031 is just visible to the naked eye of an astronomer when her pupil diameter |
| :--- | :--- | :--- | :--- | is 7 mm .

Suggest whether she can observe WASP-82 using a telescope with an objective diameter of 60 mm .
Support your answer with a calculation.
$\qquad$
$\qquad$
$\qquad$
$\qquad$

| $\mathbf{0}$ | $\mathbf{2} .5$ | $\mathbf{5}$ CCDs are often connected to telescopes. |
| :--- | :--- | :--- |

Explain two reasons why this improves the ability of astronomers to observe dim stars.

1
$\qquad$
$\qquad$
$\qquad$
2 $\qquad$
$\qquad$
$\qquad$

| $\mathbf{0}$ | $\mathbf{3}$. | $\mathbf{1}$ |
| :--- | :--- | :--- |

$\qquad$
$\qquad$
$\qquad$

| $\mathbf{0}$ | $\mathbf{3} .2$ | $\mathbf{2}$ | In 2019, astronomers linked several radio telescopes to produce a single telescope |
| :--- | :--- | :--- | :--- | called the EHT. The resolution of the EHT is the same as the resolution that a telescope with an aperture equal to the diameter of the Earth could achieve.

Table 2 shows data about the EHT and the Hubble telescope.
Table 2

|  | Aperture | Operating wavelength |
| :--- | :---: | :---: |
| EHT | $1.3 \times 10^{7} \mathrm{~m}$ | 1.3 mm |
| Hubble | 2.4 m | 410 nm |

Galaxy M87 is $5.3 \times 10^{7}$ light years from Earth. The supermassive black hole at the centre of M87 has a mass $6.5 \times 10^{9}$ times the mass of the Sun.

The radius of the event horizon is $R$.
The astronomers propose to use either the EHT or the Hubble telescope to observe stars whose distance from the centre of the black hole is less than $1000 R$.

Discuss, with calculations, which telescope is more suitable for this observation.


| 0 | 3 | 3 | A star is orbiting the black hole in M87. The star is observed in the plane of its orbit. |
| :--- | :--- | :--- | :--- | :--- | The wavelength of a spectral line observed in the light emitted from the star varies between a maximum and a minimum value.

maximum value observed $=374.96 \mathrm{~nm}$ minimum value observed $=373.53 \mathrm{~nm}$

Calculate the orbital speed of the star.

| 0 | 4 | M40 A and M40 B are two stars that appear very close to each other when viewed |
| :--- | :--- | :--- | from Earth.

There are two possible reasons for this:

- they are an orbiting binary system
- they are distant from each other and only appear in the same line of sight.

In an orbiting binary system, the difference between the apparent magnitude and the absolute magnitude for each star is similar.

Table 3 shows data about these two stars.
Table 3

|  | Temperature / K | Radius of star / m | Apparent magnitude |
| :--- | :---: | :---: | :---: |
| M40 A | 6000 | $6.3 \times 10^{9}$ | 9.7 |
| M40 B | 4700 | $1.1 \times 10^{10}$ | 10.1 |

Discuss the appearance of the two stars to an astronomer on the Earth. In your answer you should:

- compare the colour of the stars
- compare the brightness of the stars
- deduce, with a calculation, whether the stars form an orbiting binary system.


| 0 | 5 | Figure 2 shows, for some galaxies, how their recession speed $v$ varies with |
| :--- | :--- | :--- | distance $d$ from the Earth.

Figure 2

$\begin{array}{llll}0 & 5 & 1 & \text { Estimate, using Figure 2, the age in seconds of the Universe. }\end{array}$
$\qquad$

| 0 | 5 | 2 |
| :--- | :--- | :--- | The estimate in Question 05.1 assumes that the Universe has expanded at a constant rate. Measurements involving type 1a supernovae that are at large distances from Earth caused astronomers to make a modification to this assumption.

State:

- the modification
- the explanation that was proposed to account for this modification.
$\qquad$
$\qquad$
$\qquad$
$\qquad$
$\qquad$
$\qquad$
$\qquad$
$\qquad$





## Copyright information

For confidentiality purposes, all acknowledgements of third-party copyright material are published in a separate booklet. This booklet is published after each live examination series and is available for free download from www.aqa.org.uk.

Permission to reproduce all copyright material has been applied for. In some cases, efforts to contact copyright-holders may have been unsuccessful and AQA will be happy to rectify any omissions of acknowledgements. If you have any queries please contact the

Copyright © 2023 AQA and its licensors. All rights reserved.


